

More is different: Particle-number-dependence in EFT

Joint seminar in Central China Normal University



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Dec. 2, 2021



Collaborators: A. Ekström, C. Forssén, G. Hagen, U. van Kolck, G. Rupak

A story

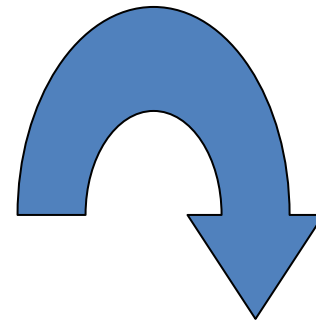




Continue driving? (change oil/fuses, add additives, perform minor fixes until it doesn't work)

Full (expensive) checks/fixes

Normally you choose...



Chiral EFT approach

to nuclear physics (Baryon-Baryon section)

History

- Weinberg (90's): Proposal of ideas.
- Bira van Kolck (94-96's): Implementations (including the Δ -full version) on NN, NNN, and EM-probe.

(Let's continue using WPC)



• Entem, Machleit, Epelbaum, Meissner, Kaiser,... (00-present): Refinements ($\chi^2 \approx 1$) and carried out to higher orders.

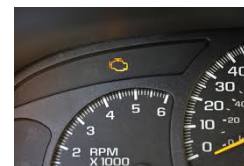
• Ekstrom et al (13-present): Refinement to fit nuclei (NNLOopt, NNLOsat, NNLOgo,... etc,)

Bayesian error analysis (on-going): Uncertainty quantifications of WPC.

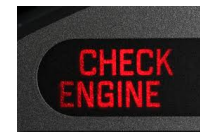
(Let's try something else)



•Kaplan, Savage, Wise (97): m_π dependence is not absorbed under WPC.



•Nogga, Timmerman, van Kolck (05): Renormalization problem (even without varying m_π).



•Birse, Valderrama, Long/Yang (06-13): New, RG-o.k. PCs.

Chiral EFT approach

to nuclear physics (Baryon-Baryon section)

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- Weinberg (90's): Proposal of ideas.
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(Let's try something else)

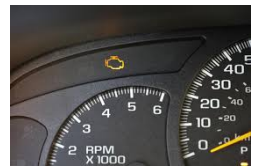
• Entanglement, Meissner, Refinement, high



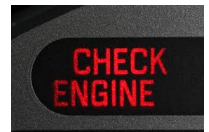
• EKS, Refinement, NNLC

Bayesian Uncertainty quantifications of WPC.

• Kaplan, Savage, Wise (97): m_π dependence is not absorbed under WPC.



• Nogga, Timmerman, van Kolck (05): Renormalization problem (even without varying m_π).



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For further information/opinions, see:

arXiv:2107.11675

and

arXiv:2111.00930

Nuclear Effective Field Theories:
Reverberations of the early days

What Can Possibly Go Wrong?

Harald. W. Griebhammer

U. van Kolck

*Université Paris-Saclay, CNRS/IN2P3, IJCLab,
91405 Orsay, France*

and

*Department of Physics, University of Arizona,
Tucson, AZ 85721, USA*

Received: date / Accepted: date

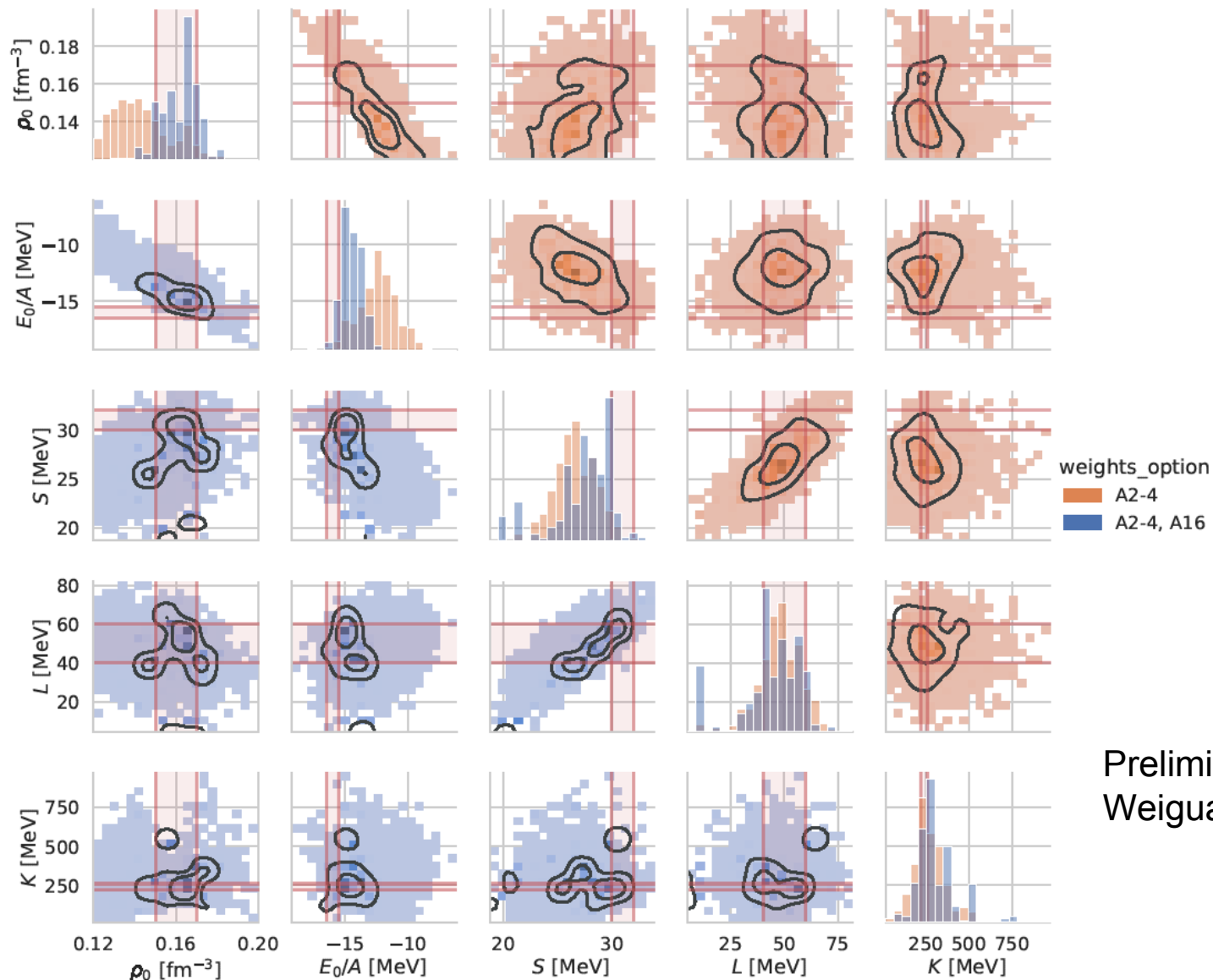
Abstract A lot.

July 27, 2021

Current status: nuclear structure

- Most ab-initio calculations adopt chiral EFT potential organized under **Weinberg power counting (WPC)**. Van Kolck, Epelbaum, Machleidt, and more.
- Good results (w.r.t. exp. data) for light systems, if low-energy constants (LECs) are renormalized at NN/NNN-level.
- **But not quite the same for ^{16}O (or heavier) → need to refit (optimize) the potential and sacrifice NN.**

Optimizing NNLO with $\Delta(1232)$



Preliminary work
Weiguang Jiang, et al.

Possibilities for improvements

WPC is wrong, one needs to organize things differently.

Need improvement

Current approach is correct, just need to include more higher-order potentials.

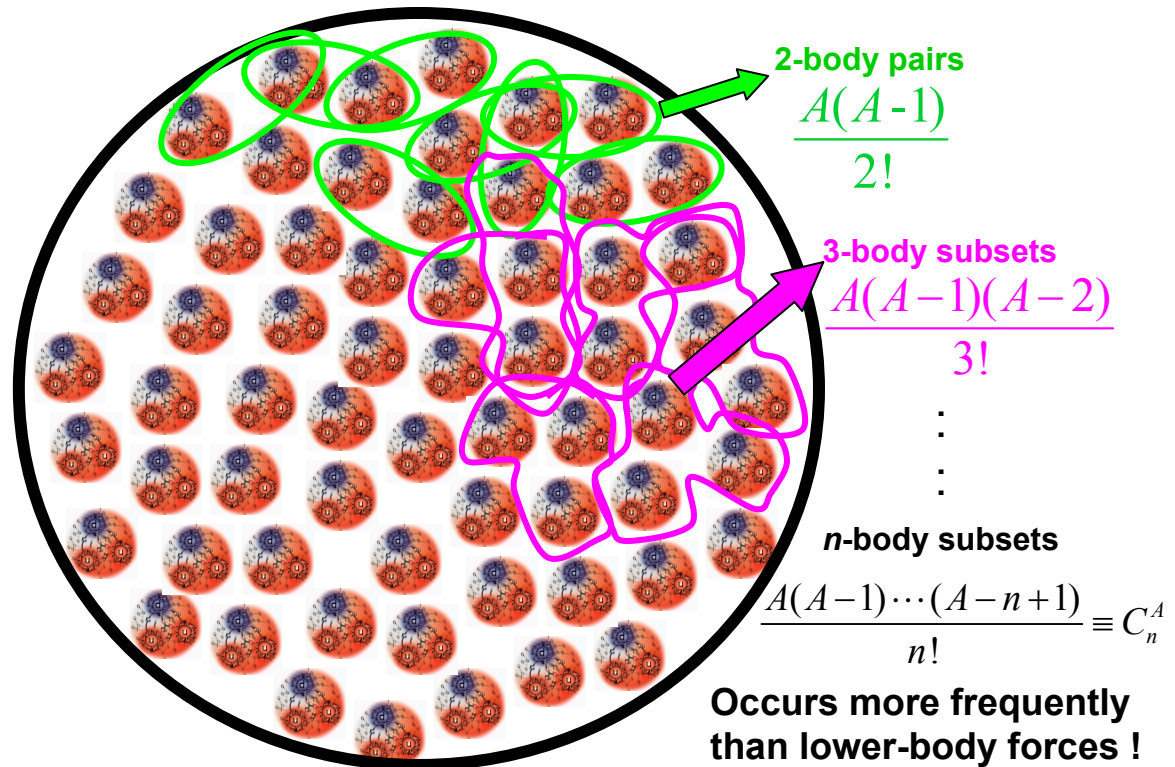
Need improvement

This work:
The importance of higher-body forces (3,4-body forces) need to grow with number of the particles in the system.
C.-J. Yang et al: [arXiv:2109.13303](https://arxiv.org/abs/2109.13303)

Many-body forces in complex systems

- Some of many-body couplings are genuine and unknown, i.e., cannot be derived from NN couplings.
- They are estimated to be weaker by naïve dimension analysis (NDA).
- However, their importance can grow in a large system.

↓ Explained in:



A	number of doublets $\frac{A(A-1)}{2}$	number of triplets $\frac{A(A-1)(A-2)}{6}$
3	3	1
4	6	4
5	10	10
6	15	20

“A choose n” enhancements

$$C_n^A = \frac{A(A-1)(A-2)\dots(A-n+1)}{n!}$$

- In a self-bound system, the above enhancement won't be fully counted. For example, an n-body subset will have nearly zero contribution if its constituents span a distance much larger than the range of the n-body forces. → density saturates, not → ∞.
- On the other hand, those small contributions could still add up to become sizable, due to the fact that there are many of them.
- Thus, the growth of n-body forces in large systems depends on multiple factors such as the **range** and the **form** of interactions, the mass of particles, etc., → **Require ab-initio calculations to know the PC.**

N-body forces in NDA

$$F_n = f_\pi^2 M_{hi}^2 \left(\frac{\nabla^2}{M_{hi}^2} \right)^s \left(\frac{N^+ N}{f_\pi^2 M_{hi}} \right)^n$$

$$f_\pi \approx 93 \text{ MeV} \quad \curvearrowright \quad 3\text{NFs}/2\text{NFs} \approx \frac{N^+ N}{f_\pi^2 M_{hi}}$$

$$M_{hi} \approx 500 - 1000 \text{ MeV}$$

$N^{(+)}$: nucleon field, $N^+ N \approx \rho$.

s: number of derivatives

*Version: Follows J. L. Friar, Few-Body Systems, 22(4):161–193 (1997) to include the 4π factor in the infrared enhancement. Otherwise n-body forces further suppressed by one chiral order $\sim 1/3$ w.r.t. 2-body forces.

Estimations

- Combine NDA and “A choose n”:

Combine both:

$$\frac{C_n^A F_n}{C_m^A F_m} = \frac{A - m}{n} \left(\frac{\rho_0}{f_\pi^2 M_{\text{hi}}} \right)^{n-m} \stackrel{\text{Approx. with nuclear saturation density}}{=} \frac{A - m}{n} \left(\frac{142(\text{MeV})}{M_{\text{hi}}} \right)^{n-m}$$

~ 1

NN and NNN becomes the same important starting from $A=13-26$ ($M_{\text{hi}}=500-1000$ MeV)

*NNN and NNNN becomes the same important starting from $A=17-34$.

*5⁺-body force is more suppressed ($s \geq 1$), only equal to NNNN after $A > 500$.

*Might be weaker due to the previous page's argument.

Let's start from light systems: ignore 3NFs first

For ${}^3\text{H}$, ${}^3\text{He}$, ${}^4\text{He}$, use no-core shell model (NCSM).

For ${}^{16}\text{O}$ and nuclear matter, use coupled-cluster method (CC).

We want to test RG also, so
adopt modified Weinberg counting
(MWPC)

Modified Power counting

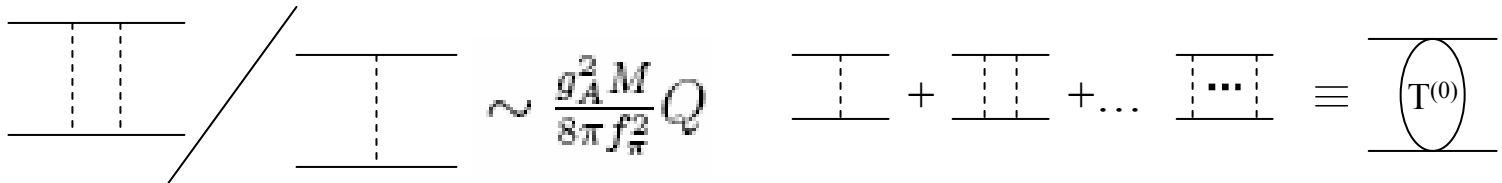
(Decided by RG)

Long & Yang, (2010-2013).

LO: Still iterate to all order (at least for $l < 2$).

Reason: van Kolck, Bedaque,... etc.

Thus, $O(Q^0)$:

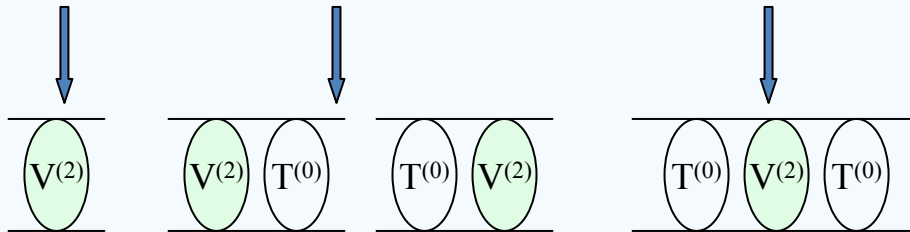


$$\text{[Series of diagrams]} \sim \frac{g_A^2 M}{8\pi f_\pi^2} Q \text{ [Single diagram]} \quad \text{[Series of diagrams]} \equiv \text{[Diagram with } T^{(0)} \text{ loop]}$$

Start at NLO, do perturbation. ($T = T^{(0)} + T^{(1)} + T^{(2)} + T^{(3)} + \dots$)

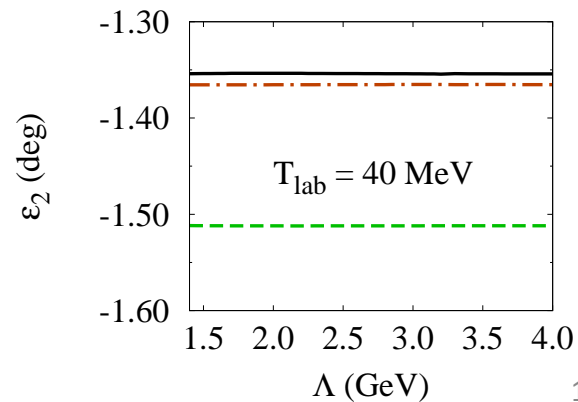
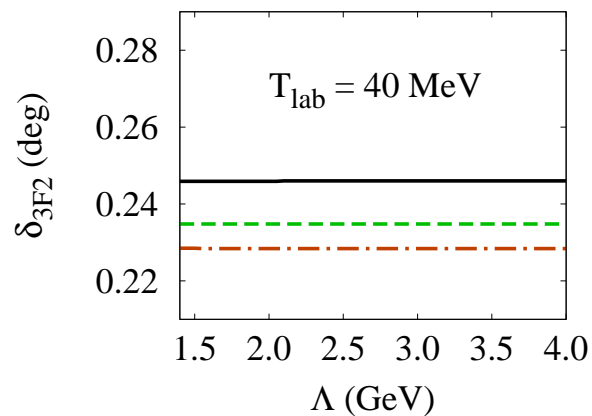
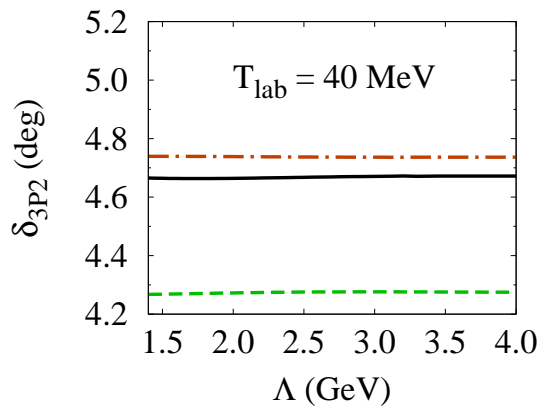
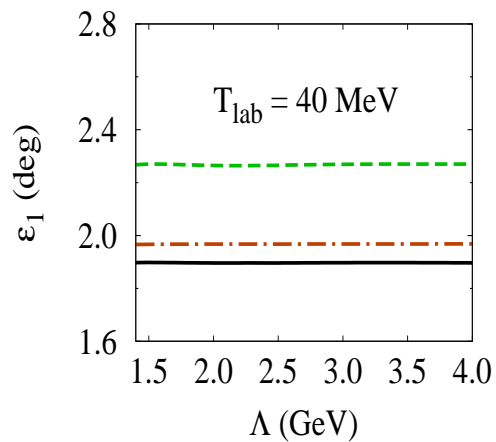
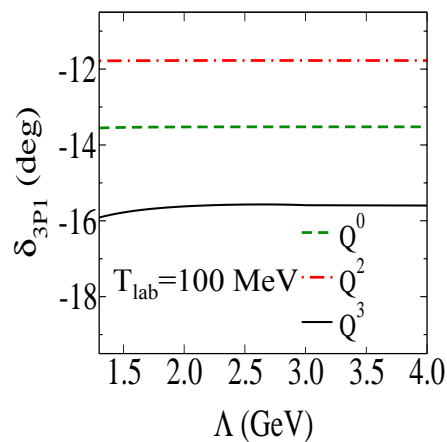
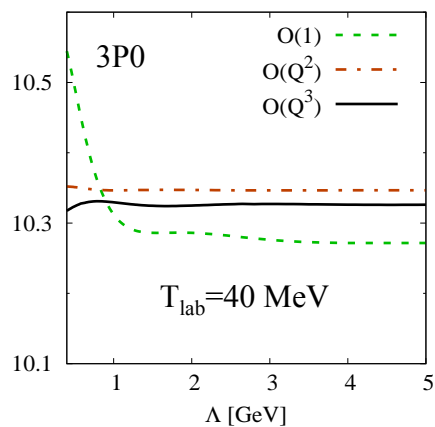
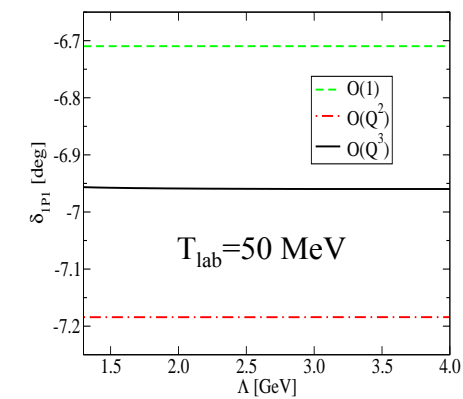
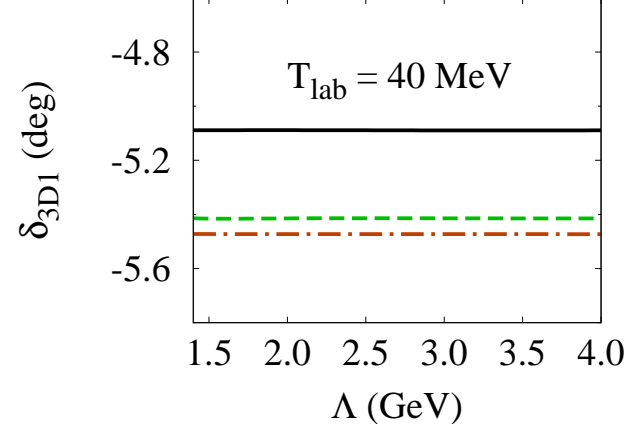
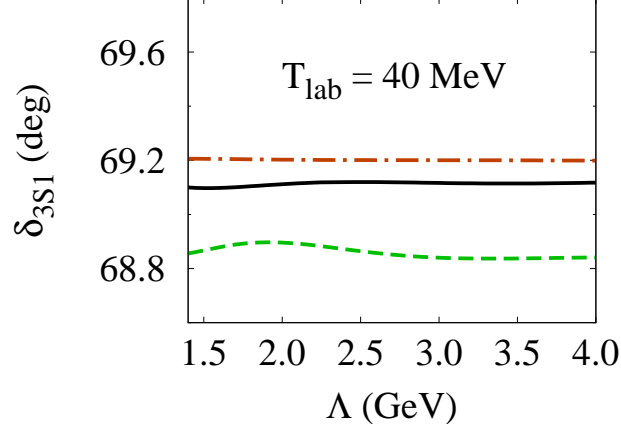
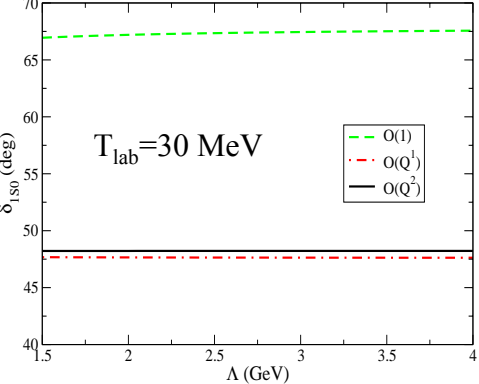
If $V^{(1)}$ is absent:

$$T^{(2)} = V^{(2)} + 2V^{(2)}GT^{(0)} + T^{(0)}GV^{(2)}GT^{(0)}.$$

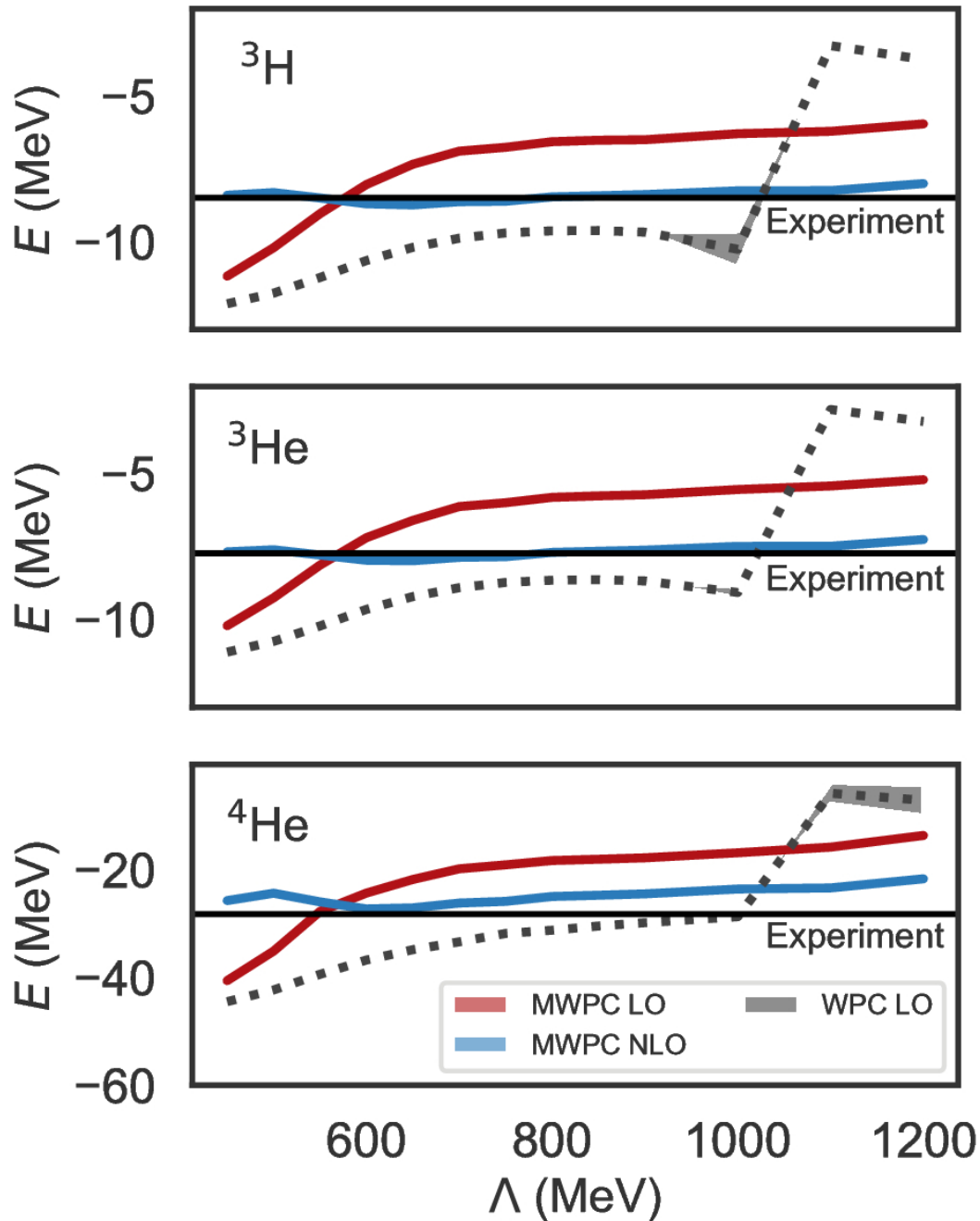


$$G \equiv \frac{2M_N}{\pi} \int_0^\Lambda \frac{p^2 dp}{p_0^2 - p^2 + i\epsilon}$$

$$T^{(3)} = V^{(3)} + 2V^{(3)}GT^{(0)} + T^{(0)}GV^{(3)}GT^{(0)}.$$



Few-body results



Conclusion:

New power counting (MWPC) works, at least for $A \leq 4$ systems.

C.-J. Yang, A. Ekstrom, C. Forssen, G. Hagen,
<https://arxiv.org/abs/2011.11584>

For A up to 3 see also:
 Nogga et al, PRC 72 (2005), 054006
 Song et al, PRC 96 (2017), 024002.

Wrong ^{16}O pole

The same NN interaction generates ^{16}O with the *wrong pole structure* (not stable w.r.t. 4α decay) at LO. Also, deformed state becomes deeper than spherical state.

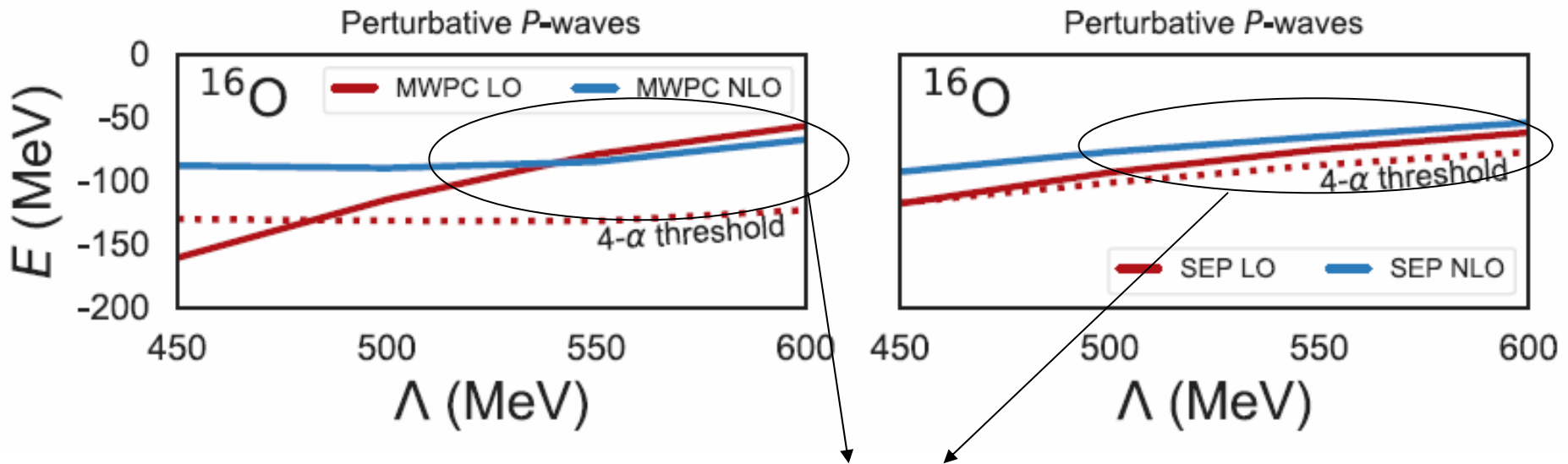
Same thing for PC improved with auxiliary dibaryon fields, Weinberg counting and pionless EFT.



M. S. Sánchez, C.-J. Yang, Bingwei Long, U. van Kolck, Phys.Rev. C97 (2018) no.2, 024001.

In fact, nobody got ^{16}O right at LO yet!

^{16}O results (LO, NN only)



^{16}O non-physical !

MWPC:

At LO, Nogga, Timmerman, van Kolck PC

(*Phys.Rev.C* 72 (2005) 054006)

NLO, plus Long & Yang PC

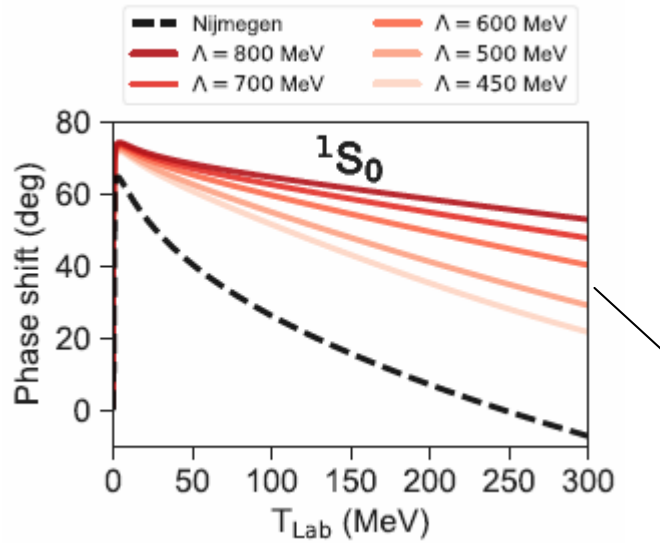
(*Phys.Rev.C* 86 (2012) 024001)

SEP: NN 1s0 adopts dibaryon field

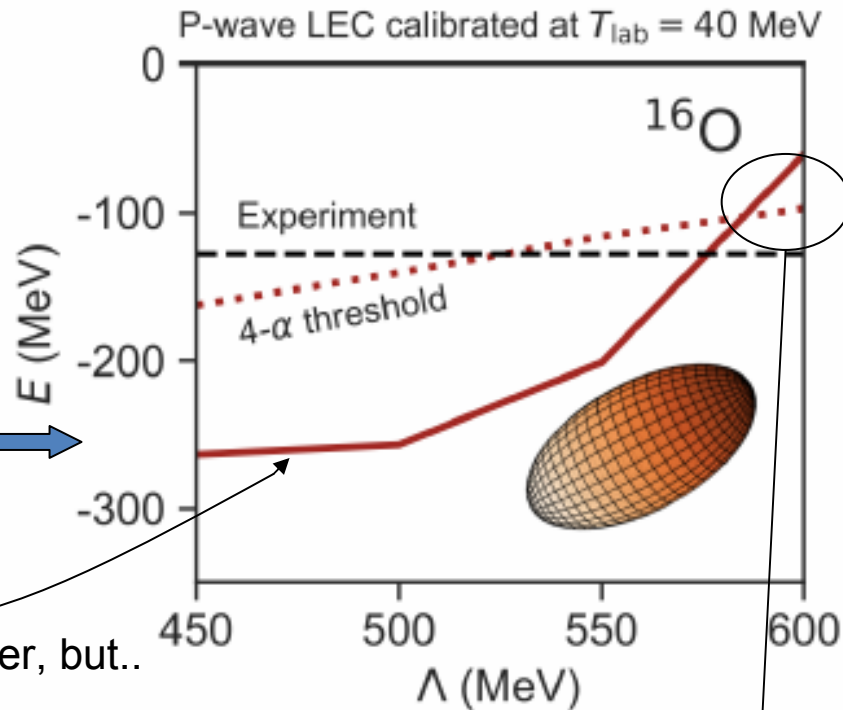
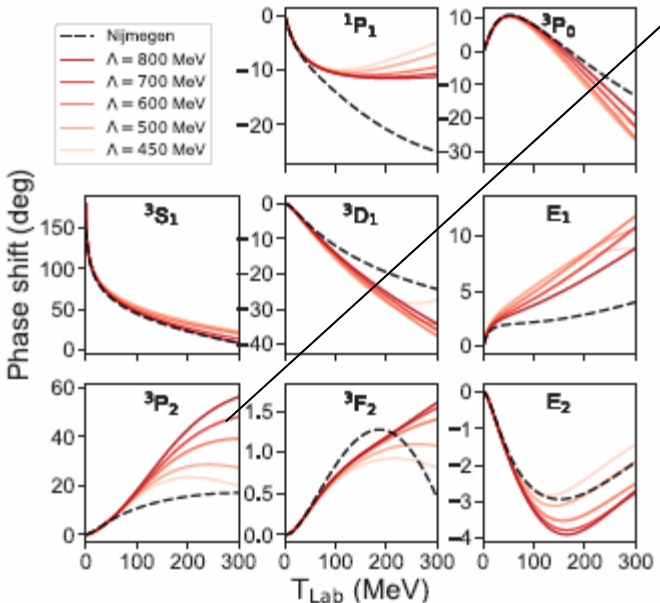
(*Phys.Rev.C* 97 (2018) 2, 024001)

Perturbative P-waves: PC by S. Wu & B. Long (*Phys.Rev.C* 99 (2019) 2, 024003)

Further fine-tune (LO, NN only)



over-attractive



Make ^{16}O deeper, but..

^{16}O is still non-physical
(deformed and unbound)!

- We have exhausted all possibilities (dibaryon, perturbative P-waves, different fitting of LECs) we could think of in the NN sector.
- Maybe need NNN force at LO.

Estimations

- Combining NDA and “A choose n”:

Combine both:

$$\frac{C_n^A F_n}{C_m^A F_m} = \frac{A - m}{n} \left(\frac{\rho_0}{f_\pi^2 M_{\text{hi}}} \right)^{n-m} \approx \frac{A - m}{n} \left(\frac{142(\text{MeV})}{M_{\text{hi}}} \right)^{n-m}$$

Approx. with nuclear saturation density

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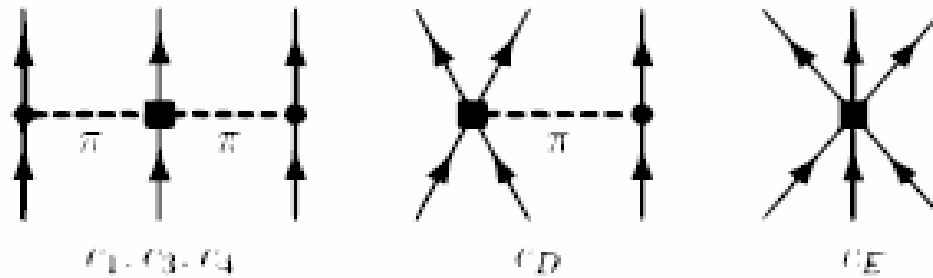
*Might be even weaker due to the previous page's argument.

**NN and NNN becomes the same
important starting from A=13-26 !**

^{16}O has $A=16!$

3-body force at LO for ^{16}O

What to include?



NNLO in WPC

All LECs are fixed by NN, π N

This diagram contributes repulsively
In ^{16}O or larger nuclei. For $A \leq 4$ it's
also weakly repulsive.

C_D , C_E are free parameters. Renormalization
needed.

It turns out, adding only one of them will
destroys RG.

(Values $C_{1,3,4}$ from Siemens, et al, PLB 770, 27 (2017))

Boundary condition is important

Including only 1 diagram won't work (Long and Bira, Annal. phys. 323 (2008) 1304)

Λ (MeV)	c_E	$E(^3\text{H})$ (MeV)	$E(^4\text{He})$ (MeV)
450	-0.40	-8.48	-24.37
500	-0.205	-8.48	-24.96
550	-0.066	-8.48	-25.69
600	0.0534	-8.48	-26.7
650	0.158	-8.48	-28.09

$C_D = 0$, C_E fixed to ^3H binding

^4He **diverges w.r.t. cutoff!**

Need both C_D and C_E

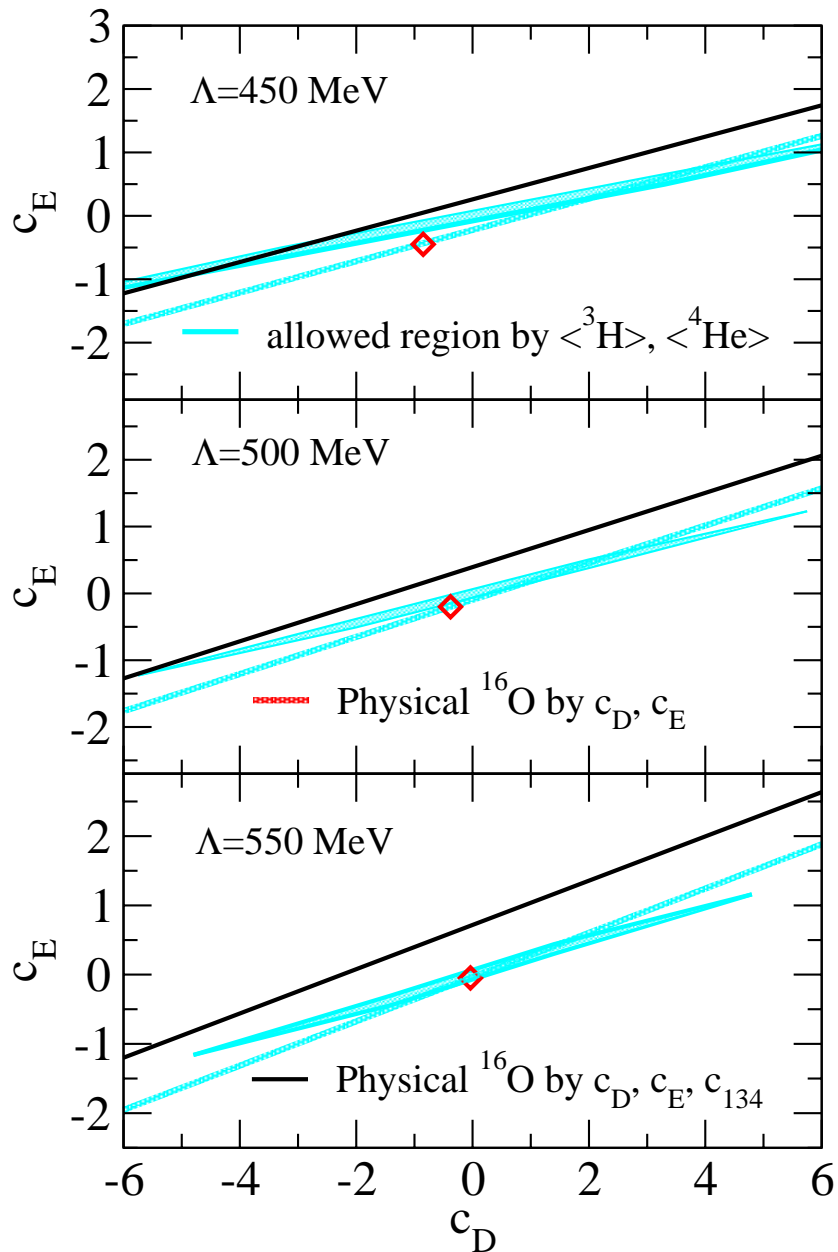
how to fix them?

Linear correlated C_D, C_E

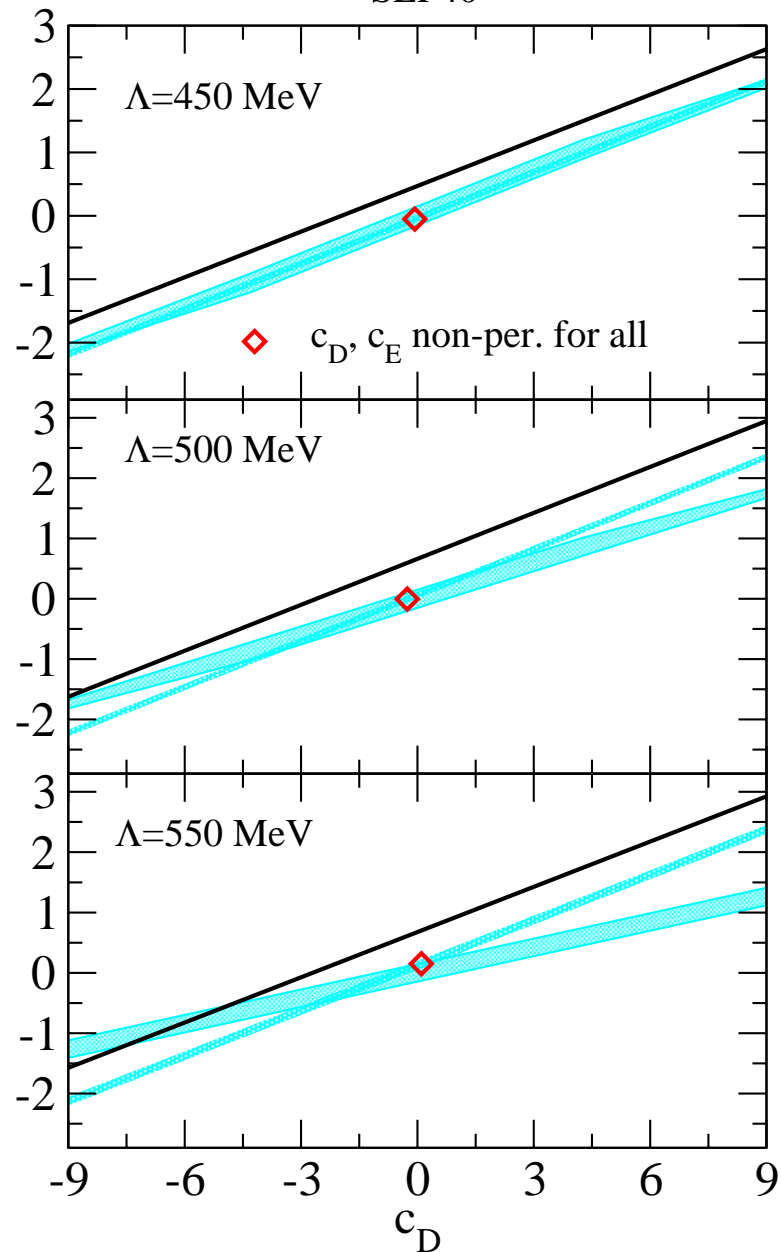
- C_D, C_E, C_{134} should enter at **higher order** (probably at NNLO or later) for **$A \leq 4$ nuclei**.
- But they will enter **non-pertur. at LO for ^{16}O** .

=>Require $|\langle \phi_{\text{LO},3\text{or}4} | C_D + C_E | \phi_{\text{LO},3\text{or}4} \rangle|$ to be smaller than 1/6 of the $^3\text{H}, ^4\text{He}$ binding (need to satisfy in both $A=3$ & 4 cases). **→ Denoted as the blue region in next page.**

MWPC40

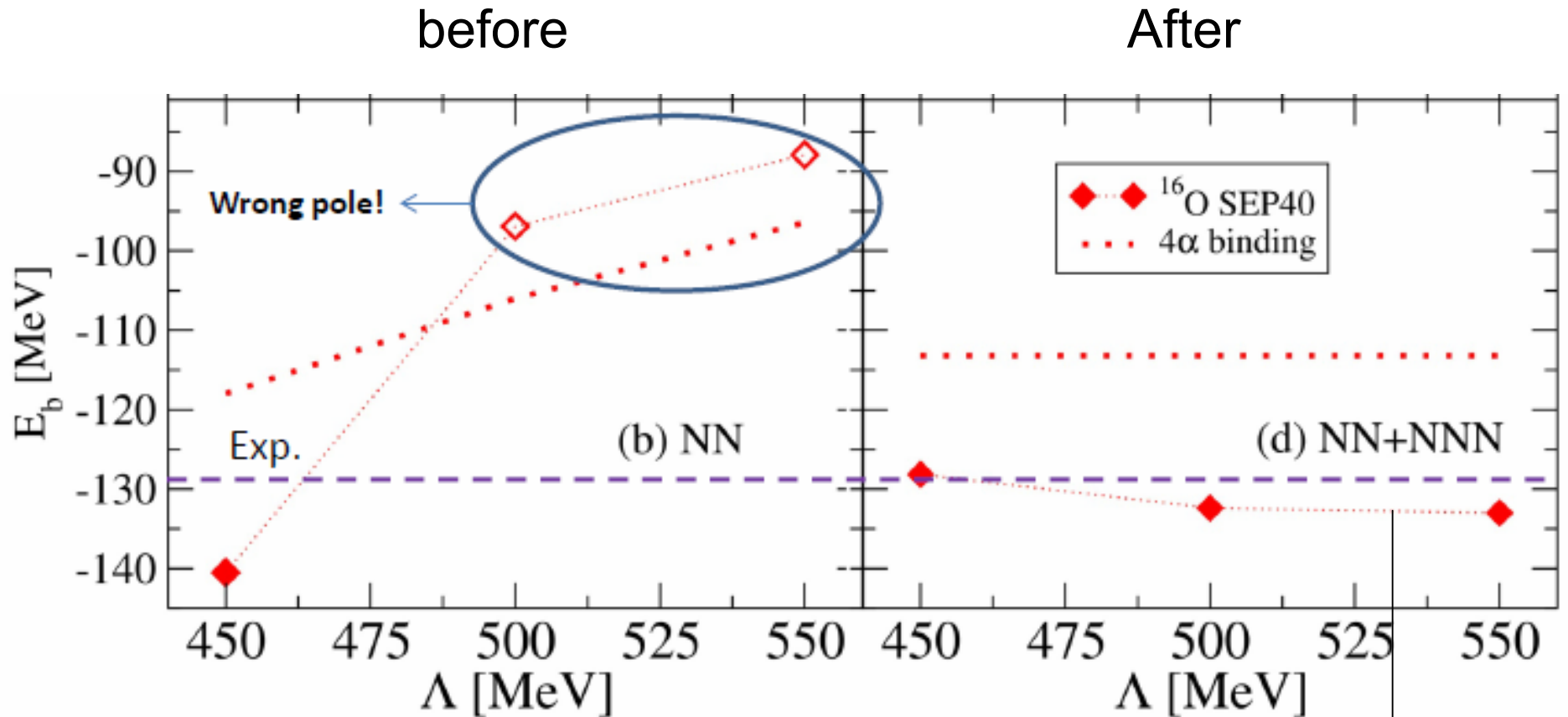


SEP40



Red and black lines are the (c_D, c_E) and (c_D, c_E, c_{134}) reproduce ${}^{16}\text{O} = -128 \pm 10\%$.

Using the overlap of the **red** and **blue** lines, i.e., (C_D, C_E) fixed by $\langle {}^3\text{H} \rangle$ and $\langle {}^4\text{He} \rangle$, we got:

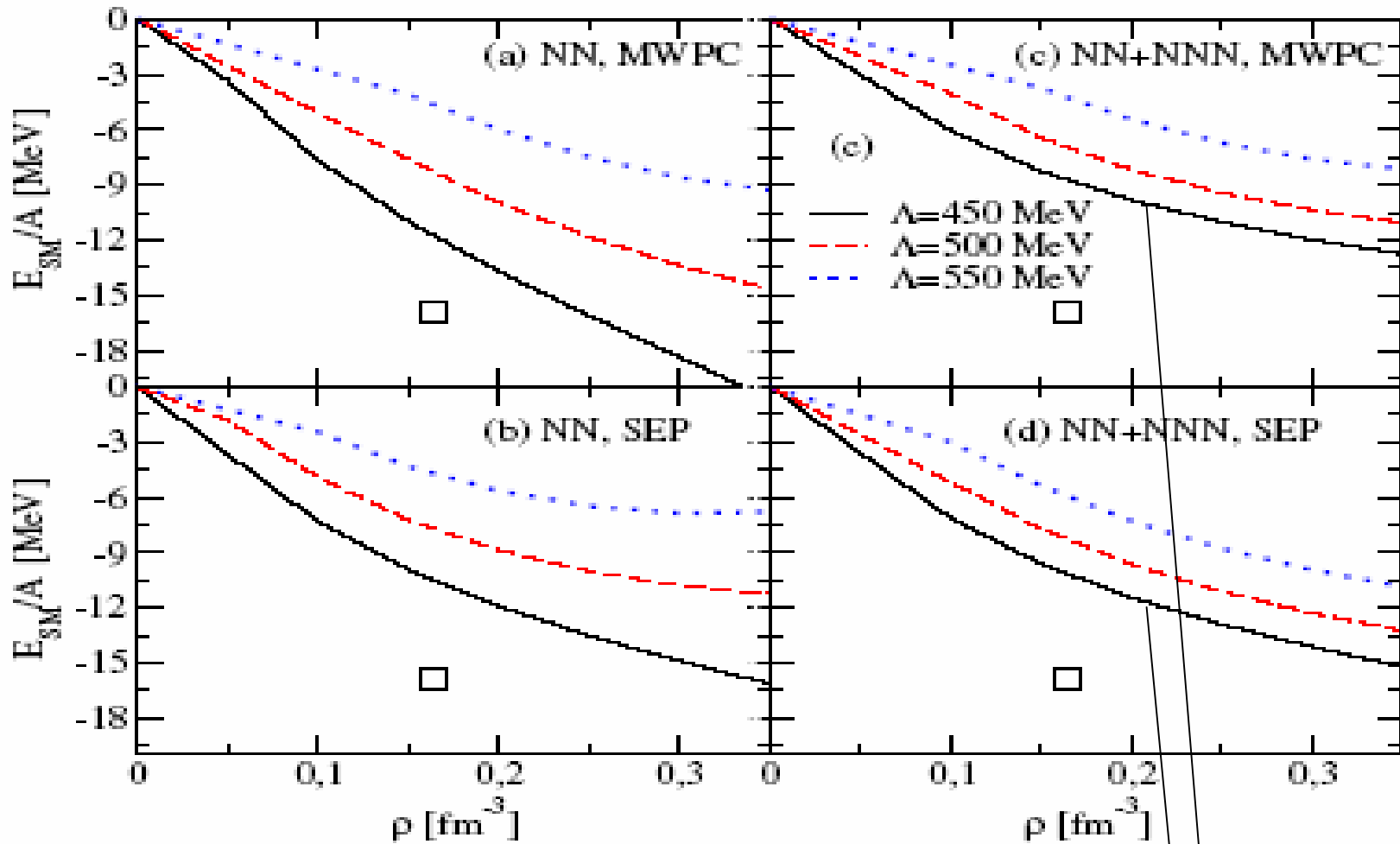


Problem solved! ${}^{16}\text{O}$ great already at LO!

This suggests:
3NFs are LO at least for $A \geq 16$

NNNN at LO?

No-go test by nuclear matter (EoS)



Conclusion: NN+NNN seems **no enough** !

Comments on current status of ab-initio approaches

- With the advance of ab-initio techniques, today the biggest **uncertainty** of nuclear structure calculations comes from the **input potential**. => **One needs to improve this first.**

Without A-dependent PC

- State-of-the-art calculations includes at least V_{NNLO} in WPC, i.e., where 3NFs first appears.
- Fitting LECs at few-body sector → Bad saturation properties.
- Fitting LECs overall → Sacrifices NN and few-body sector.

Our study suggests

- **Maybe we really need to include the A-dependent power counting! Where the orders 3NFs and 4NFs enter dep. on A.**
(more flexible in some sense, as their LECs don't need to accommodate every nuclei at the same time.)

Summary

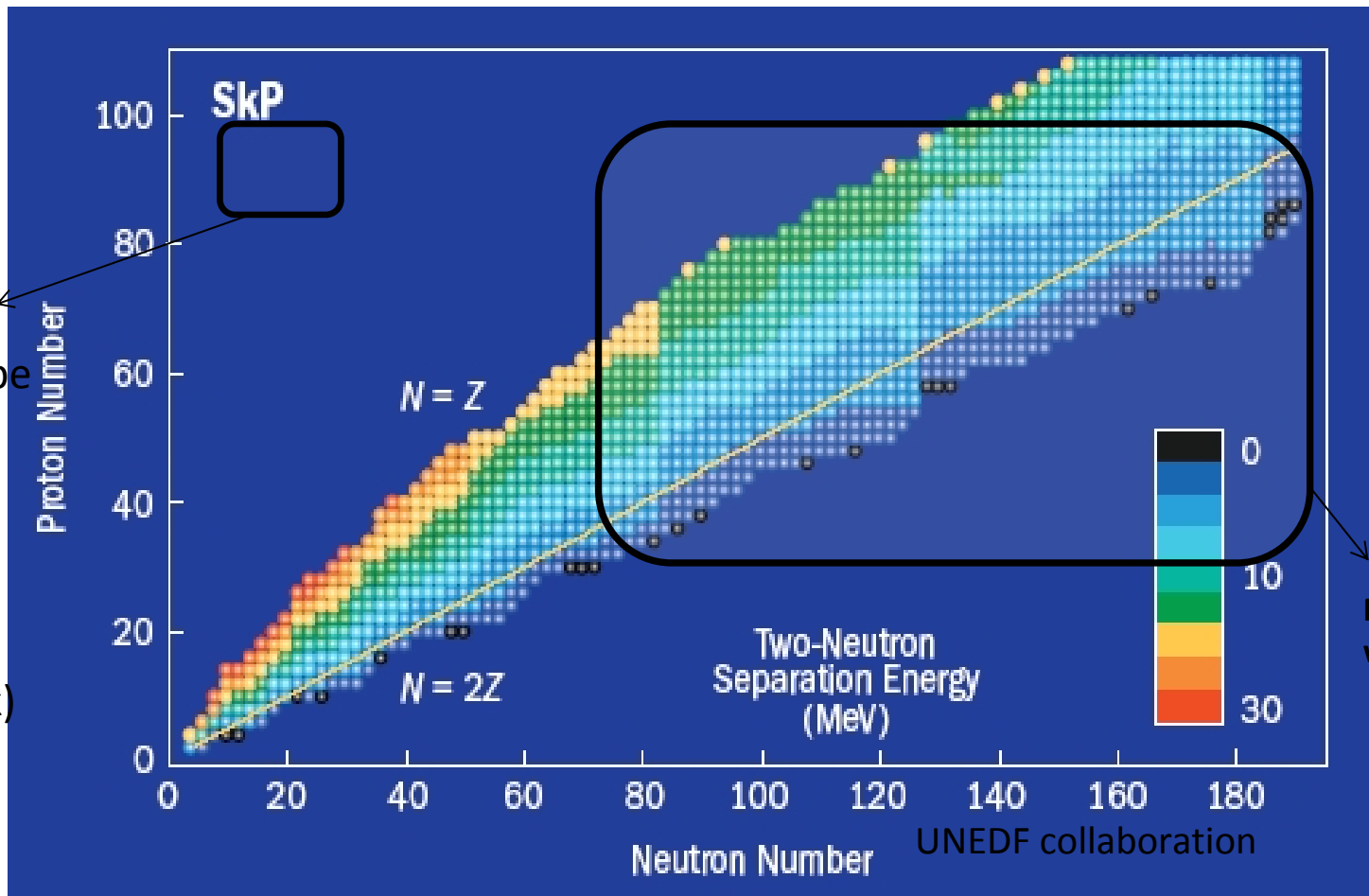
- Starting from $A=16$ (or lower), we need to promote 3-body force (c_D, c_E) at LO in chiral EFT.
- For saturation properties to be reasonably described, a further promotion of 4-body forces to LO is likely.

Solutions:

- In short term, might play with Λ and regulator form to avoid 4-body forces ($c_D(\Lambda), c_E(\Lambda)$... go through zero).
- In the long term, need to further develop ab-initio codes/technique to include up to 4-body forces.

Any other way (to get heavier nuclei)?

Mean field with Skyrme-type



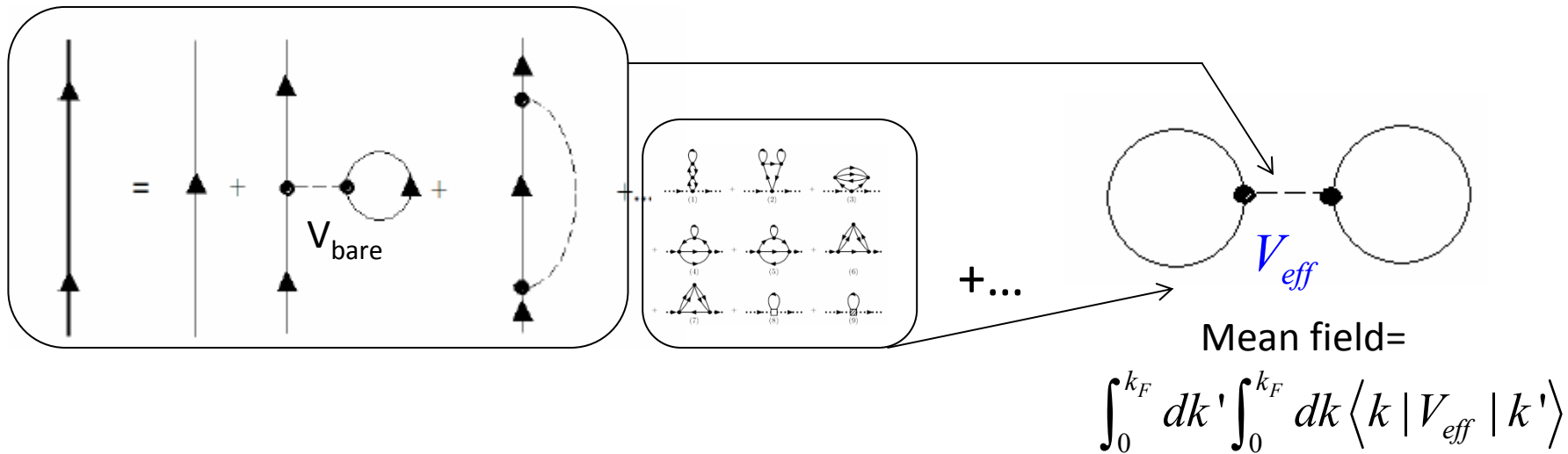
Skyrme-type interaction works o.k. (able to do the fitting in EDF framework)

No way to get with ab-initio!

Need to think about other expansion (than on NN d.o.f.).

From bare interaction $\rightarrow V_{\text{eff}}$

Analogue: from QCD to V_{NN}



$$V_{\text{eff}} = \underbrace{t_0(1 + x_0 P_\sigma)}_{S\text{-wave } O(0)} + \frac{1}{2} \underbrace{t_1(1 + x_1 P_\sigma)(k'^2 + k^2)}_{S\text{-wave } O(q^2)} + \underbrace{t_2(1 + x_2 P_\sigma) \mathbf{k}' \cdot \mathbf{k}}_{p\text{-wave } O(q^2)}$$

$$+ \frac{1}{6} \underbrace{t_3(1 + x_3 P_\sigma) \rho^\alpha}_{s\text{-wave, higher body force}}$$

$$P_\sigma = \frac{1}{2} (1 + \sigma_1 \cdot \sigma_2)$$

I.e., CTs in pionless EFT

Has many-body force to begin with!

By the way, this (MF EDF) is a pure perturbative approach (i.e., you *do not* iterate V_{eff} at all).

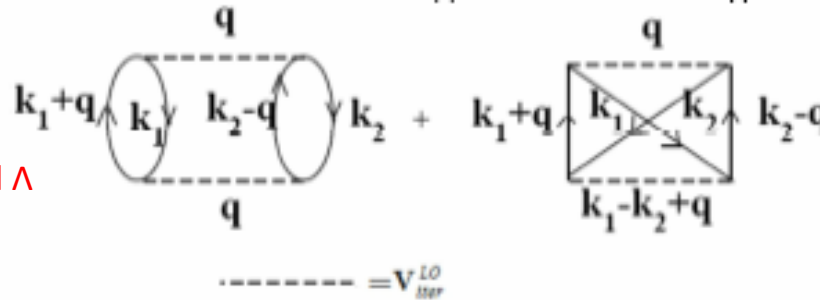
One would like to add some EFT features to it.

First step, need an NLO correction to MF.

Between ab-initio and EDF

Diagrams of vgv:

diverge, need Λ



MBPT(ab-initio)

V defined in **vacuum**

Improvable: $v+vgv+\dots$

Exact when include all diagrams (**iterated to all order**).

Pros:

First principle calculations with clear foundation.

Issues:

Calculations too costly (model-space is large).
Could be very sensitive to detail of V .



Our approach

V re-defined at each order

Approach by: $v+vgv+\dots$, but we want to stop at certain order.

Advantages: Hopefully have advantages from both EDF & MBPT.

EDF

V defined in **medium** (mean field)

Improvable: $\langle nlm|V|nlm\rangle$, i.e., V itself, or need multi-reference.

Pros: Simple and effective

Issues:

Missing correlations beyond MF,
 V is a model !

Interpretation

LO=MF in EDF

Subleading: included via Perturbation theory

A novel approach toward finite nuclei: between *ab initio* and energy–density–functional theories

C.J. Yang,^{1,2} W.G. Jiang,¹ S. Burrello,³ and M. Grasso⁴

¹*Department of Physics, Chalmers University of Technology, SE-412 96, Göteborg, Sweden*

²*Nuclear Physics Institute of the Czech Academy of Sciences, 25069 Řež, Czech Republic*

³*Institut für Kernphysik, Technische Universität Darmstadt, 64289 Darmstadt, Germany*

⁴*Université Paris-Saclay, CNRS/IN2P3, IJCLab, 91405 Orsay, France*

We propose a novel idea to construct an effective interaction under energy–density–functional (EDF) theories which is adaptive to the enlargement of the model space. Guided by effective field theory principles, iterations of interactions as well as enlargements of the model space through particle-hole excitations are carried out for infinite nuclear matter and selected closed-shell nuclei (⁴He, ¹⁶O and ⁴⁰Ca) up to next-to-leading order. Our approach provides a new way for handling nuclear matter and finite nuclei within the same scheme, with advantages from both EDF and *ab initio* approaches.

MBI

V defini

Improvable: $v+vgv+\dots$

Exact when include all diagrams (iterated to all order).

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[arXiv:2110.01959](https://arxiv.org/abs/2110.01959)



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field)

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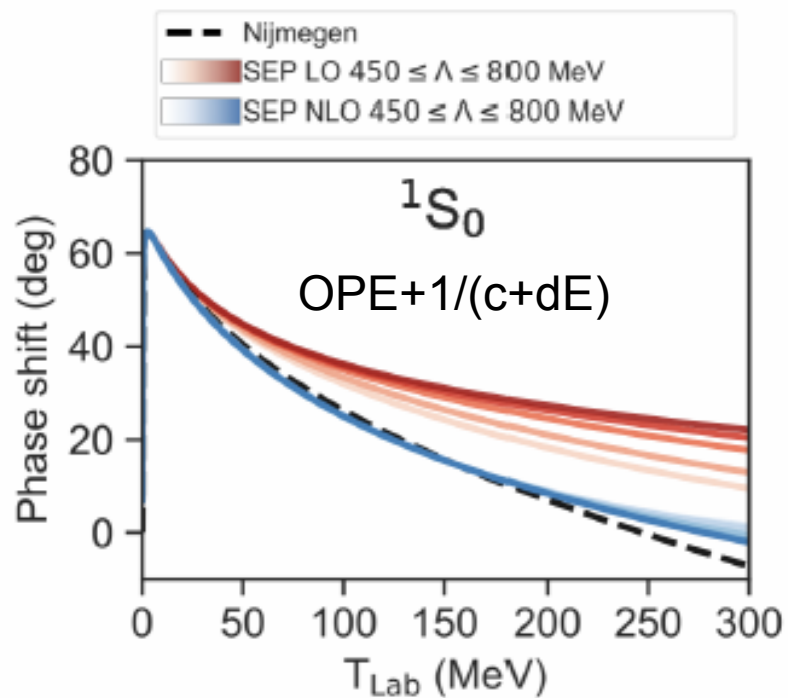
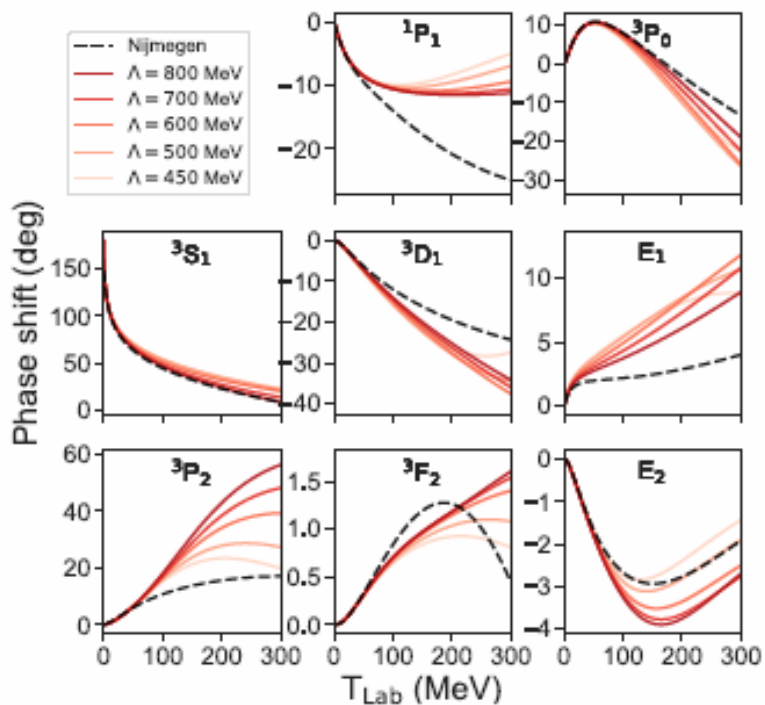
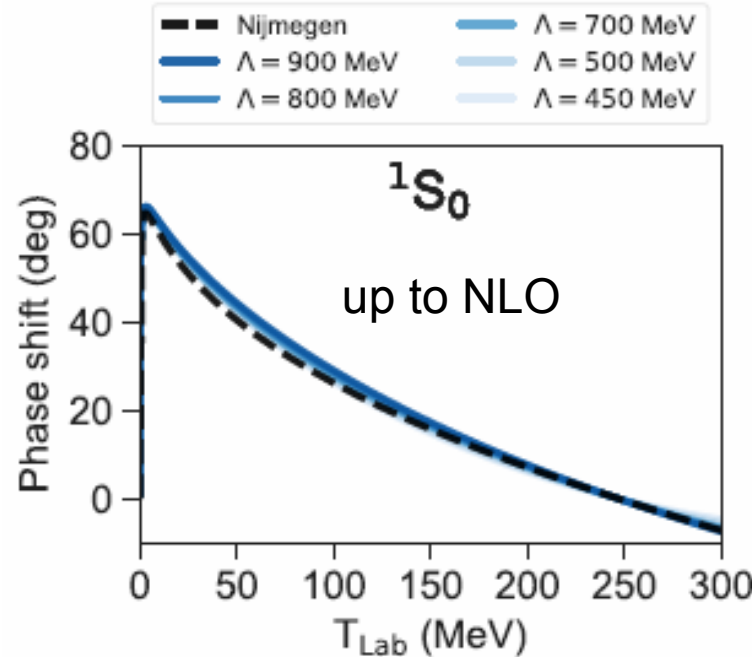
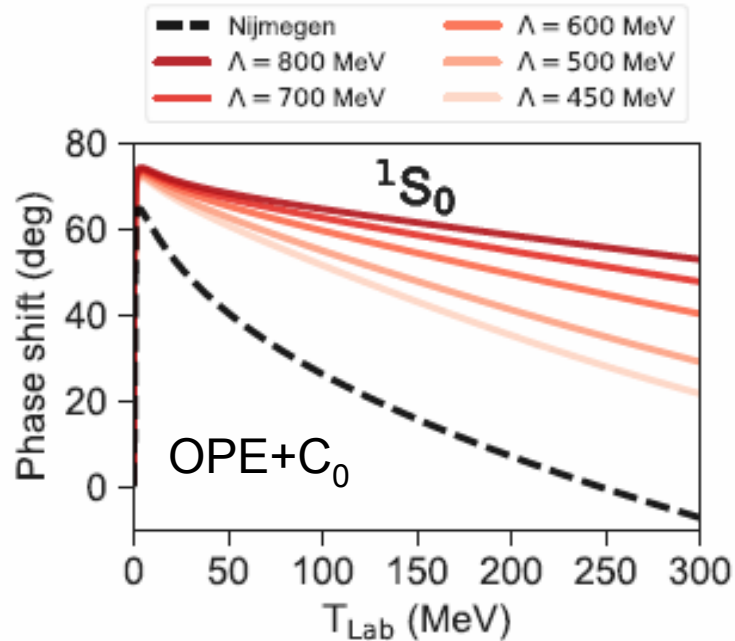
Outlook

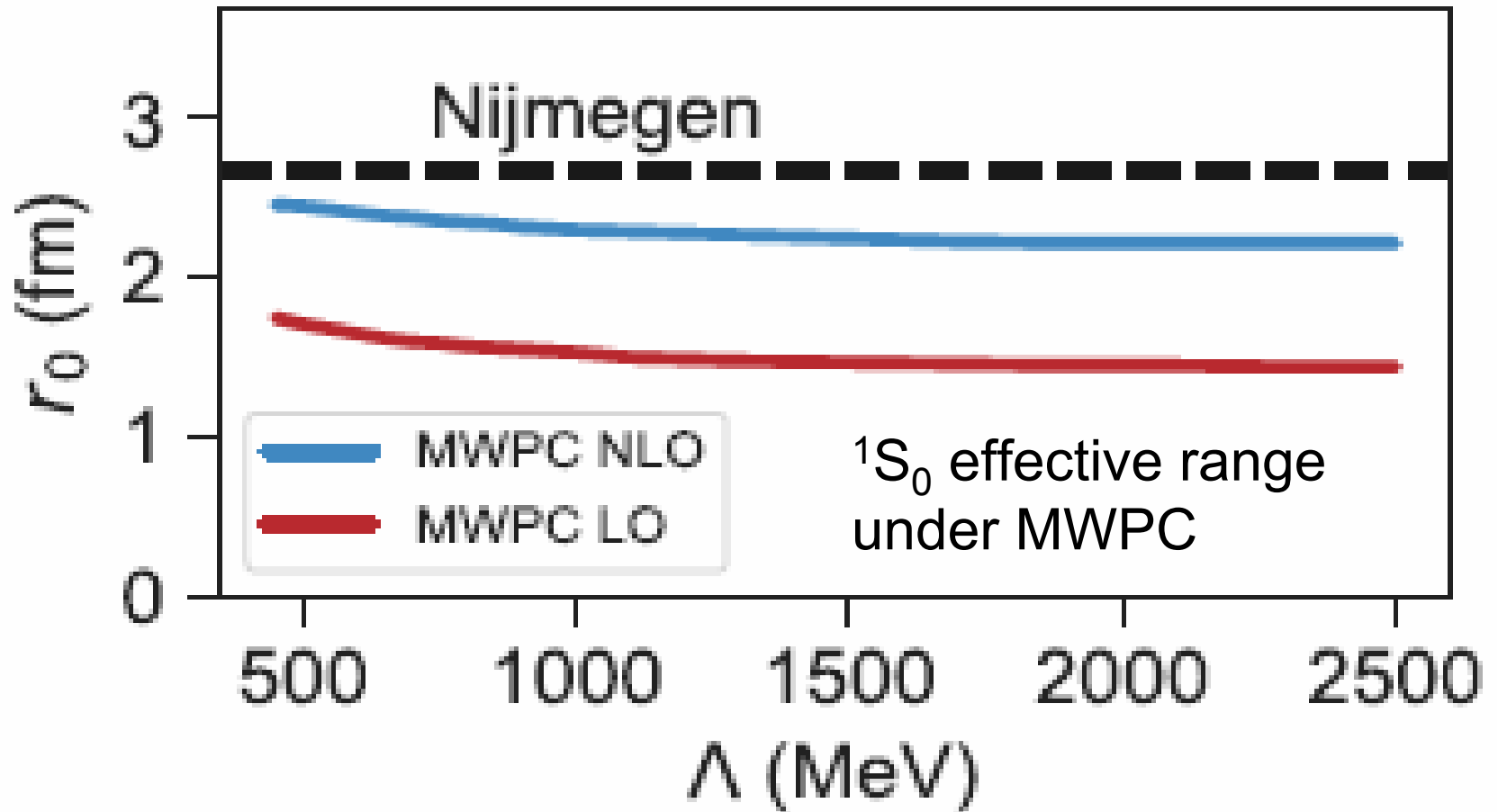
- We are at the **beginning** of a whole new, exciting paradigm!
- Lots of following works (NLO, NNLO, $\Delta(1232)$, PC for nuclei away from closed shell, core of the neutron star, etc.) to be done.

Thank you!

Potential	Lambda	B3LO	B4LO	<cD>_3	<cE>_3	<cD>_4	<cE>_4	<c1>_3	<c3>_3	<c4>_3	<c1>_4	<c3>_4	<c4>_4
MWPC40	450	-11.17156	-40.61872	-1.4	-3.66	-8.5	25.7	-0.0102	1.816	-2.25	1.065	25.175	-21.375
MWPC40	500	-10.19628	-35.12708	-0.71	1.97	-5.98	13.9	-0.0884	1.26	-2.49	0.768	25.24	-24.85
MWPC40	550	-9.07575	-29.02803	-0.0825	0.45	-0.61	2.7	-0.166	-0.192	-2.01	0.0387	11.24	-17.84

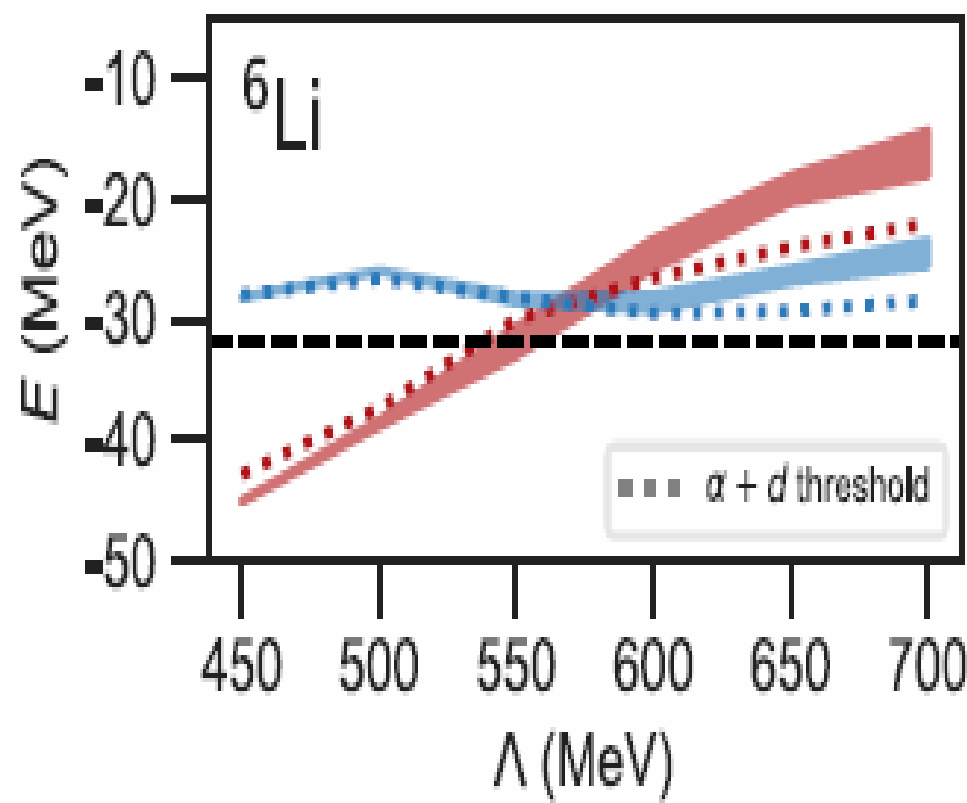
Potential	Lambda	B3LO	B4LO	<cD>_3	<cE>_3	<cD>_4	<cE>_4	<c1>_3	<c3>_3	<c4>_3	<c1>_4	<c3>_4	<c4>_4
SEP40	450	-8.705024	-29.522573	-0.064	0.23	-0.518	1.59	-0.0484	0.439	-1.067	0.313	9.85	-10.21
SEP40	500	-8.197143	-26.527799	-0.235	0.026	-1.74	0.16	-0.106	-0.289	-0.989	-0.0736	4.85	-8.83
SEP40	550	-7.727850	-24.116944	0.067	-0.77	0.425	-4.53	-0.155	-1.36	-0.667	-0.46	-3.25	-5.1



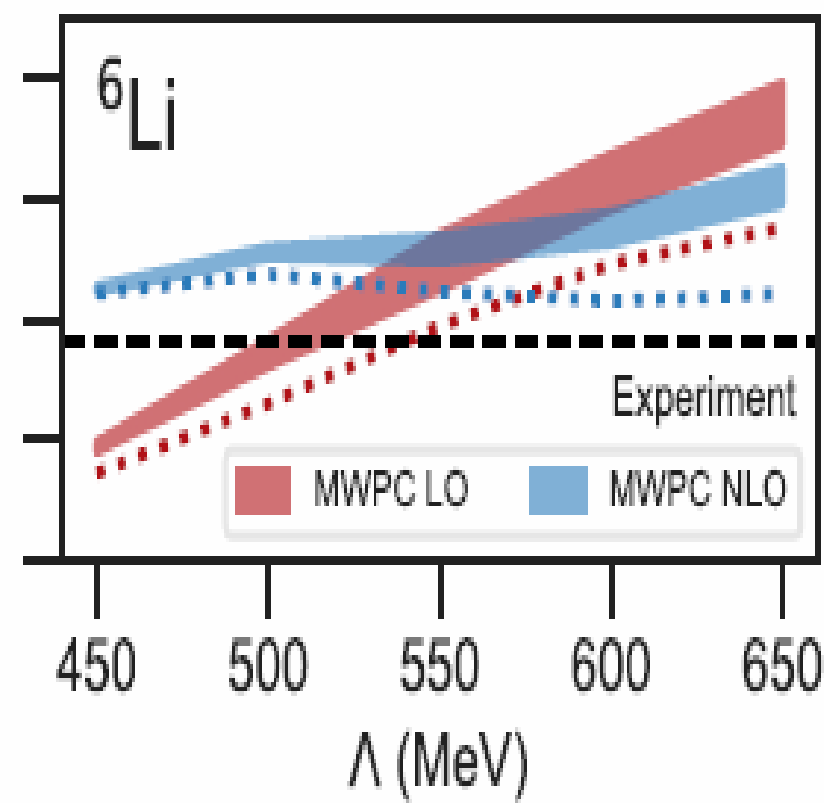


*SEP (dibaryon field) has r_0 fitted to data already at LO.

P-wave LEC calibrated at $T_{\text{lab}} = 40$ MeV



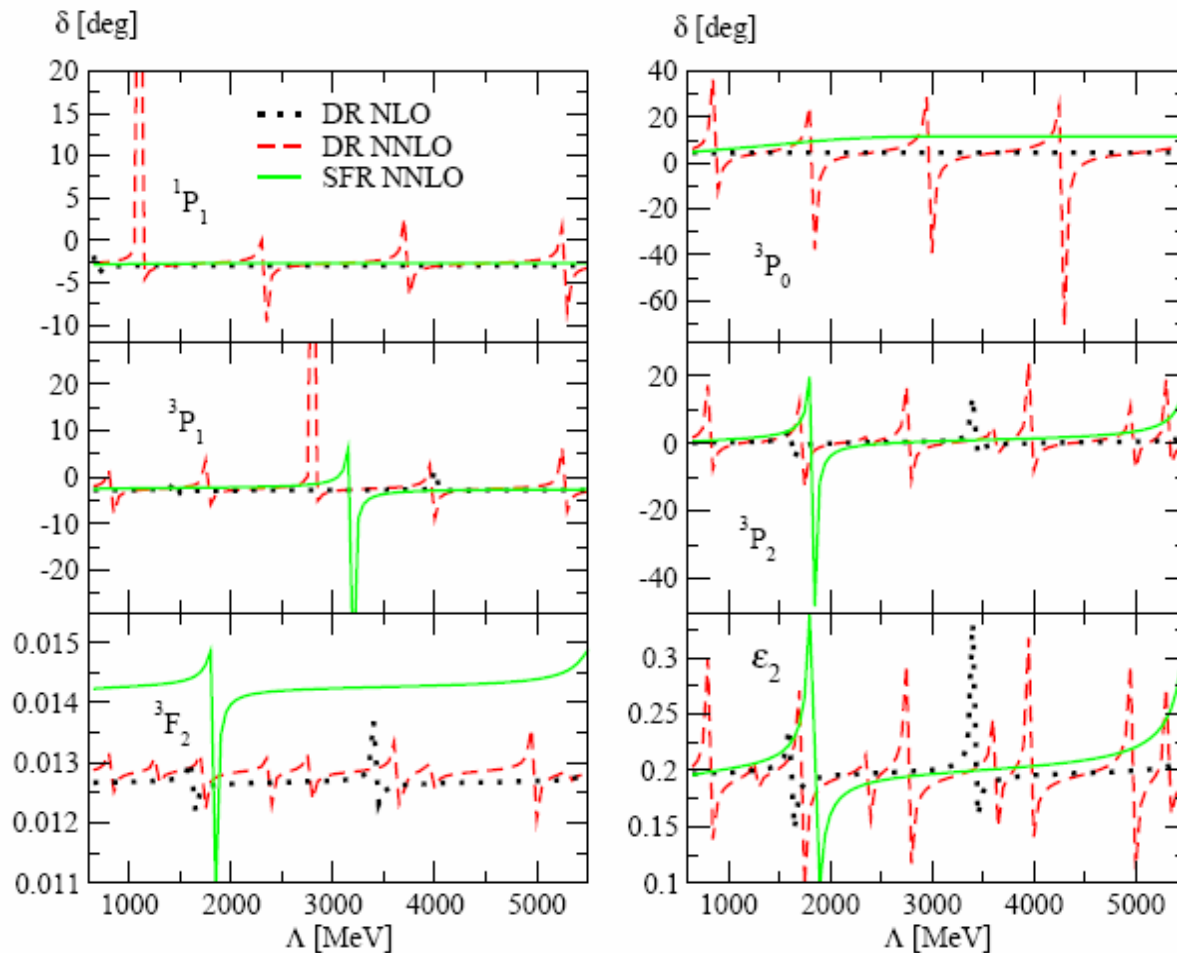
P-wave LEC calibrated at $T_{\text{lab}} = 200$ MeV



Problems of WPC

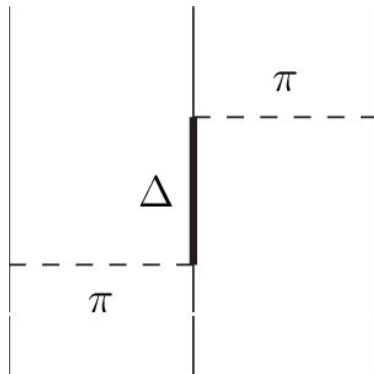
WPC is wrong at LO ! (Nogga, Timmermans, van Kolck, *PRC* 72 (2005) 054006)

• Beyond LO: (Yang, Elster, Phillips (2008-2010))



$\Delta(1232)$ NNN?

Fujita-Miyazawa



Appear at NLO, i.e., one order before c_D, c_E, c_{134} .

So far we include it by the “*resonant saturated*” way[1,2], but then it appears to be repulsive for ^{16}O , at least up to the highest cutoff (550 MeV) we can test.

[1] Epelbaum, et al, Nucl. Phys. A 806, 65 (2008).

[2] Ekstrom, et al, PRC 97, 024332 (2018).

EFT-like approaches

Select a model space

(since we cannot do calculations in (or even actually know the structure of) the entire Hilbert space)

Build/Derive the interactions in that model space

(normally utilize properties of symmetry)

Solve for any observable desired

Compare to data (check your assumptions)

Renormalization group (RG)

Select a model space

(since we cannot do calculations in (or even actually know the structure of) the entire Hilbert space)

Build/Derive the interactions in that model space

Solve for specified observable desired

Vary your model space (mostly enlarge → more shouldn't hurt)

↓
Compare to data (check your assumptions)

Minimum requirement of EFT[©]

For more details: see, e.g., publication list: 1

Select a model space

(since we cannot do calculations in (or even actually know the structure of) the entire Hilbert space)

Renormalize the interactions in that model space

Describe observable within the validity of EFT

Vary your model space (mostly enlarge → more shouldn't hurt)



Compare to data (check your assumptions)